

Possible options for XEUS secondary microcalorimeter array and TES microcalorimeter development in Japan

Kazuhisa Mitsuda ISAS, JAXA

for the meeting on A European calorimeter for XEUS October 26-27, 2004





Possible options

- Primary high resolution spectrometer
 - $\sim 32 \times 32 = 1024$ pixels, 250µm × 250µm/pixel
 - resolution $\Delta E \sim 2 \text{ eV}$ in 0.3 10 keV band
- Secondary high resolution spectrometer ~32x32 = 1024 pixels
 - 1. Wide field of view with degraded energy resolution

 $\Delta E \sim 10 \text{ eV but} > 1 \text{mm X 1mm / pixel}$ still too narrow (?)

- 2. Low energy band with better energy resolution $\Delta E < 1 \text{ eV}$ in 0.3 - 1 keV
- 3. Wide-band (hard X-ray) spectrometer
 - $\Delta E < 40 \text{ eV}$ in 0.3 100 keV, ~1mm X 1mm /pixel
- 4. Combination of the above options



Low energy options

- With thin X-ray absorbers, good energy resolution can be achieved with TES
 - Energy resolution of TES microcalorimeters

$$\Delta E \propto \sqrt{E_{sat}T} \le 1 \text{eV} \text{ for } E_{sat} = 1 \text{keV}, T = 100 \text{mK}$$

- Science
 - Emission lines from high red-shift objects
 - Absorption line forest in quasar spectra





Wide energy-band options



Multi-layer coated XEUS super mirror

2000 cm² up to ~ 40 keV
can be extend to higher energy (?)



- Soft Gamma-ray emission lines
- e.g. ⁴⁴Ti from SNR
- Cyclotron absorption lines

not so many interesting objects (?)





Our 1st trial of wide-band TES calorimeter



 $\Delta E=160 \text{ eV} @ 60 \text{keV}$ was obtained with our very first device

Absorber: Tin 0.9 mm x 0.9 mm 300 μm

Development of TES calorimeter array for the NeXT mission







TES calorimeter array

Small-absorber (180µm) calorimeter



calorimeter of the same absorber size





Low temperature signal multiplexing

demonstrated frequency-division multiplexing of two calorimeter pixels.







Iyomoto et al. 2004





Cryogenics

J-T cooler (10mW @ 1.8 K)



ADR ($T_{lowest} = 40 \text{mK}$)



Narasaki et al. 2003

